



# **Upcoming Changes to the EPA Photochemical Assessment Monitoring Stations (PAMS) Network; New Meteorological Measurement Requirement for Mixing Heights and Current Activities**

**Kevin A. Cavender  
National Program Manager  
USEPA/OAQPS**

**James Szykman  
U.S. EPA  
Office of Research and Development  
National Exposure Research Laboratory**

**Ad-hoc Mixing Layer Height Working Group  
December 6, 2016**



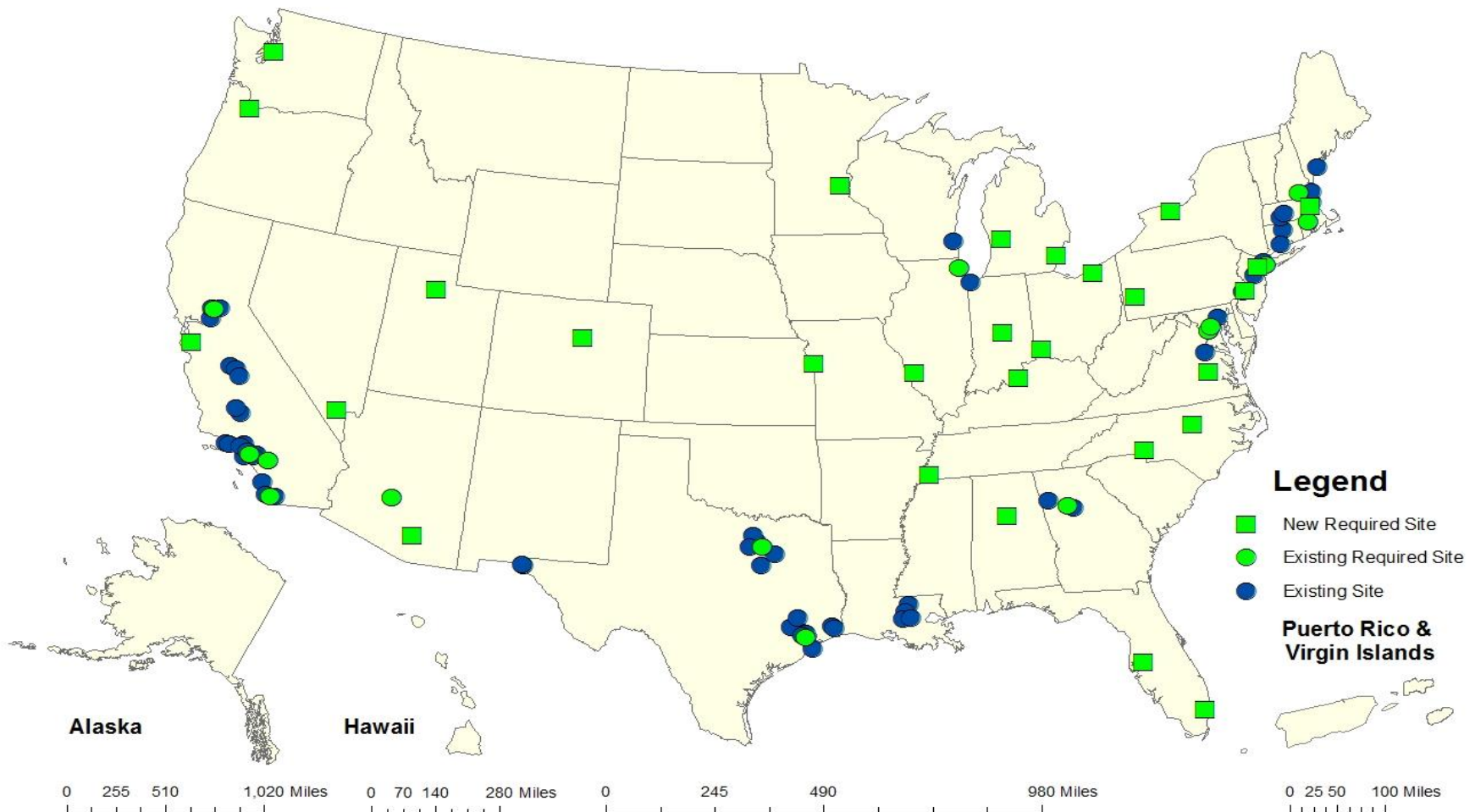
## Updates to PAMS Network Design

- Major changes to the PAMS requirements were finalized in October 2015 as part of the ozone NAAQS review
- We replaced the existing 20 year-old multi-site, enhanced ozone network design with an updated 2-part network design
  - Requiring PAMS measurements to be collocated with existing NCore sites in areas with population of 1 million or more irrespective of Ozone NAAQS attainment status
    - Results in a stable network of approximately 40 required sites with improved spatial distribution and reduced redundancy
    - Includes a waiver for historically low ozone areas
    - Includes an option to make PAMS measurements at an alternative location (e.g., an existing PAMS site) which may cross CBSA or even state boundaries
  - Require states with moderate or above ozone non-attainment areas and states in the Ozone Transport Region to develop and implement an Enhanced Monitoring Plan (EMP)
    - Provides support for flexible approaches for collecting data to understand ozone issues in new and existing high ozone areas

DRAFT – Do Not Cite Or Quote



## New and Existing PAMS Sites





## Changes to Required PAMS Measurements

- Requires hourly VOC measurements
  - Included a waiver to allow 3 8-hr canister samples in locations with low VOC concentrations and for “logistical and programmatic constraints”
- Requires 3 8-hr carbonyls samples on a 1 in 3 day schedule
  - Included an alternative to allow for continuous formaldehyde measurements
- Requires “true NO<sub>2</sub>” in addition to existing NO<sub>y</sub>
- • Requires hourly mixing height measurement (replaces “upper air measurements”)
  - Added a waiver option to allow measurements to be made at an alternative location (e.g., NOAA ASOS sites)
- Additional required PAMS meteorology measurements that are not part of the NCore requirements include atmospheric pressure, precipitation, solar radiation, and UV radiation

DRAFT – Do Not Cite Or Quote

## Summary of Flexibility in Requirements

- A number of waiver options are available to help provide flexibility
  - Waiver for low ozone concentrations (<85% of NAAQS)
  - Waiver to move location to alternative site
  - Waiver to use longer averaged VOC sampling (i.e., canisters) instead of autoGCs in some circumstances
  - Waiver to use off-site meteorology where appropriate
- EMPs are intended to provide support for flexible approaches for collecting data to understand ozone issues in new and existing high ozone areas
  - Just because a state isn't required to have an EMP doesn't mean they can't or shouldn't!



DRAFT – Do Not Cite Or Quote



## PAMS Timeline and Milestones

- PAMS plan due July 1, 2018 as part of Annual Network Plan
  - Consider moving this up to July 1, 2017 if waivers are needed!
- PAMS monitoring at NCore sites will need to start by June 1, 2019
  - Looking for some states to be early implementors and start getting equipment installed in 2017 and 2018
- EMPs submitted within two years of designations or by October 1, 2019, whichever is later







## EPA Commitments

- PAMS Funding reallocation
  - Start in 2017, and spread over multiple years
- National Procurements for autoGCs, true NO<sub>2</sub>, and ceilometers
- Guidance documents
  - Technical Assistance Document
  - Generic QAPP
  - SOPs for autoGCs, true NO<sub>2</sub>, and ceilometer
  - EMP Guidance
- National QA Program
- Training, Training, and more Training!
  - Data Validation/Reporting
  - AutoGC operation
  - Mixing height/Ceilometer





# PAMS MLH requirement and Modeling Needs

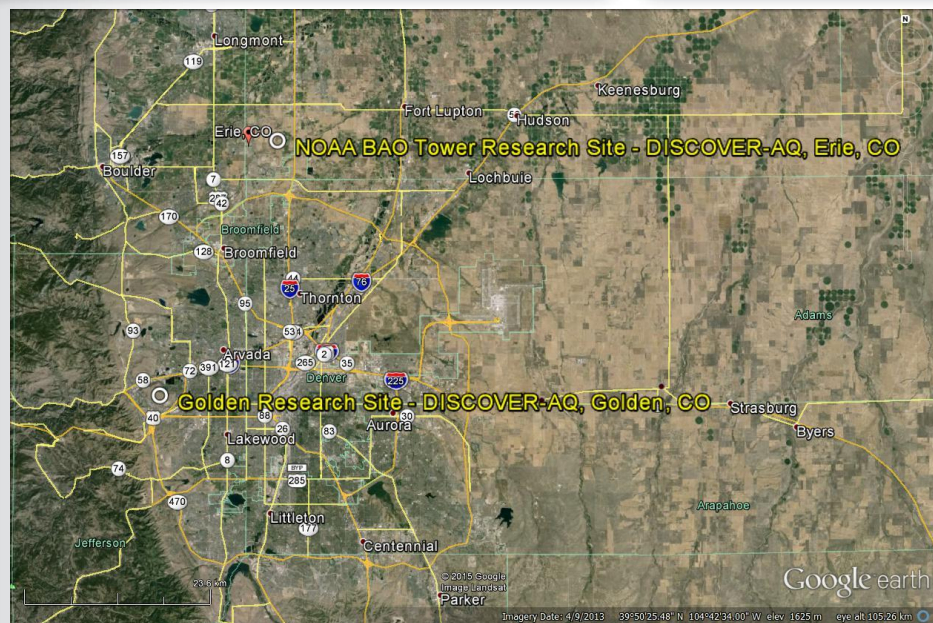
- The primary purpose for the hourly mixing heights under PAMS was driven by the SIP modeling data needs.
- Within EPA the following areas have been identified as concerns wrt ability of ceilometer/lidar measurements to provide useful data:
  - 1) Rate of PBL rise in morning and the timing of evening transition back to nocturnal layer / residual layer along with absolute MLH values.
  - 2) Measurement-to-model comparison - which parameters should be pulled from WRF (or MCIP) that best matches what the ceilometers measure,
  - 3) Models have a specific vertical layer structure which may have coarser resolution than the measurements, how can this be accounted for in assessments,
  - 4) Spatial representativeness of MLH relative to model grid cell, and relative to adjacent/nearby grid cells,
  - 5) Data gaps in data (e.g., overnight w/ lidars).

DRAFT – Do Not Cite Or Quote





## Field Evaluation locations for CL-5 I



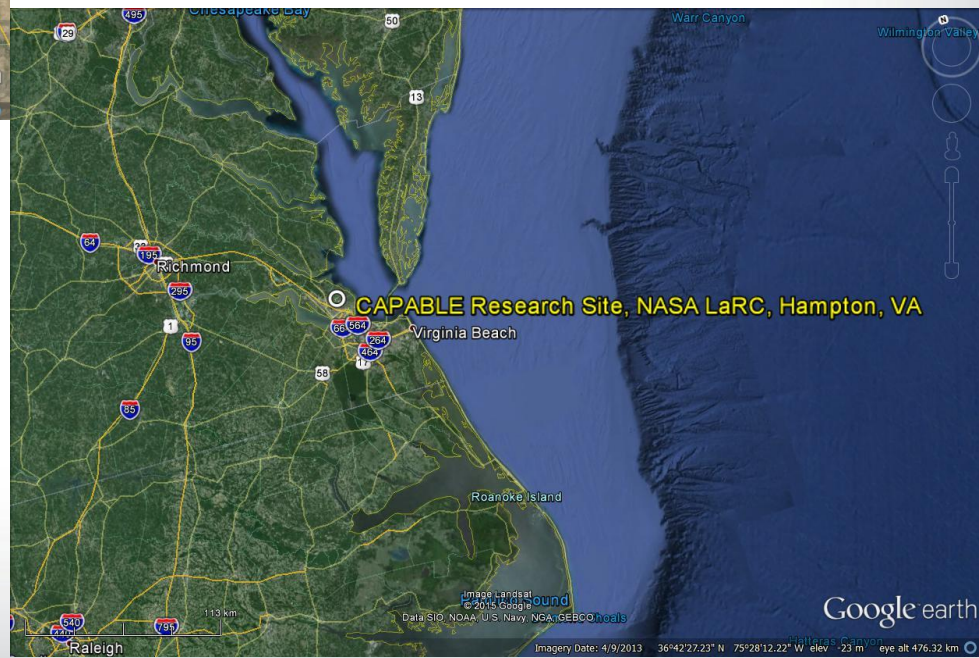
NOAA BOA Tower Site (Erie, CO)  
– start of High Plains

Golden NREL Site  
-On a mesa -intermountain site

Both sites low aerosol loading

NASA Langley (Hampton, VA)  
- Near sea level – coastal site

Low-moderate aerosol loading – with  
marine influence



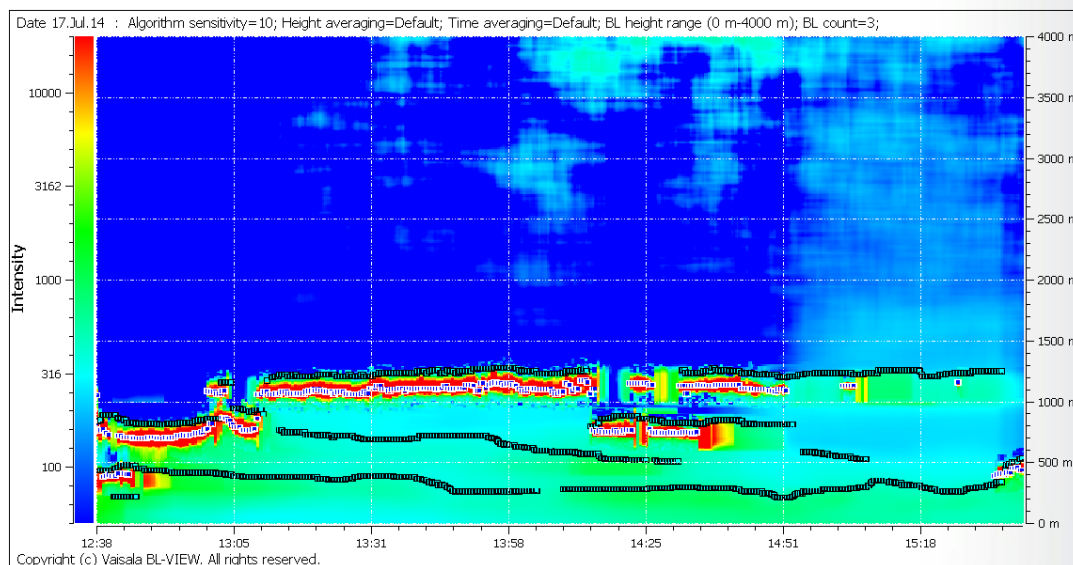


## CL-51 Comparison was conducted using the Vaisala BL-View Software for the CL-51 MLHs

### ➤ BL-View Software:

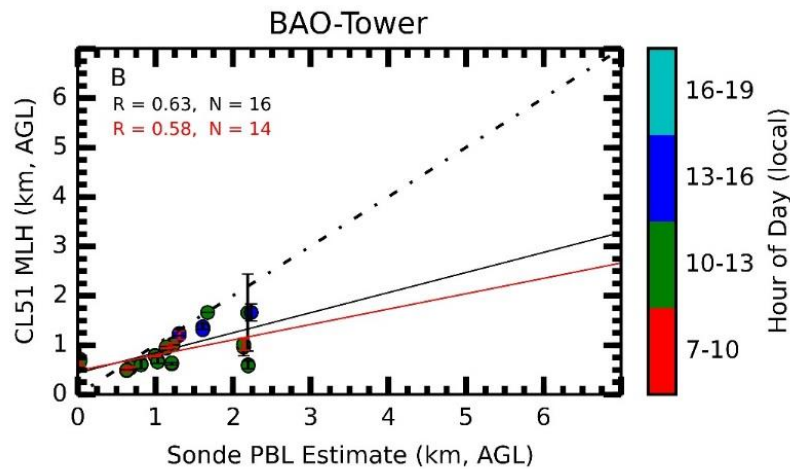
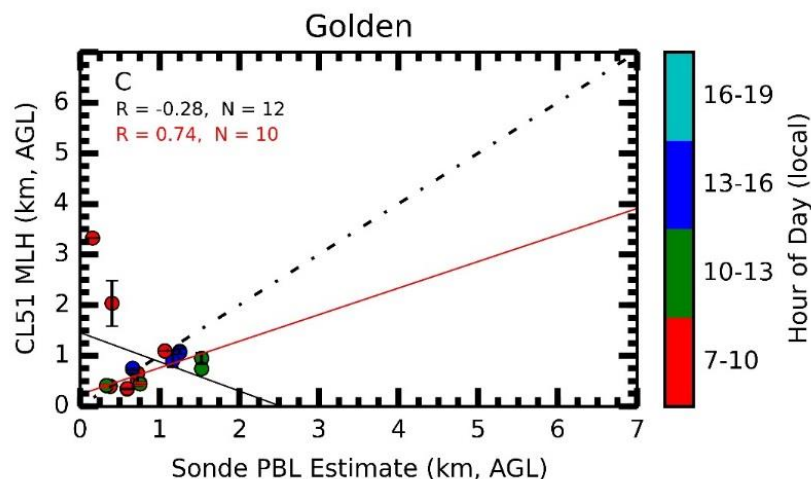
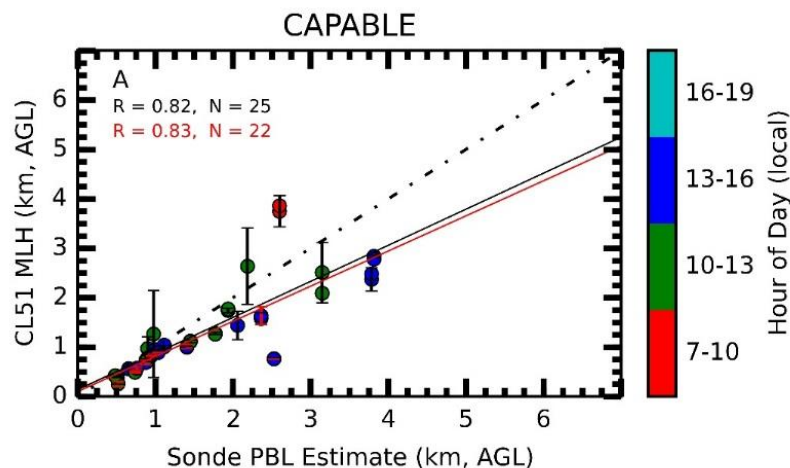
- Uses a proprietary gradient method algorithm
- Identifies up to 3 aerosol layers for consideration of MLH
- Layers assigned quality index (QI) 1 to 3; 3 highest confidence
- Use of variable time and altitude averaging

### Characteristic backscatter curtain plot generated in BL-View for 17-July 2014 Golden, CO





## CL-51 Mixing Layer Height (aerosol gradient) vs. Planetary Boundary Layer Height (thermal gradient)



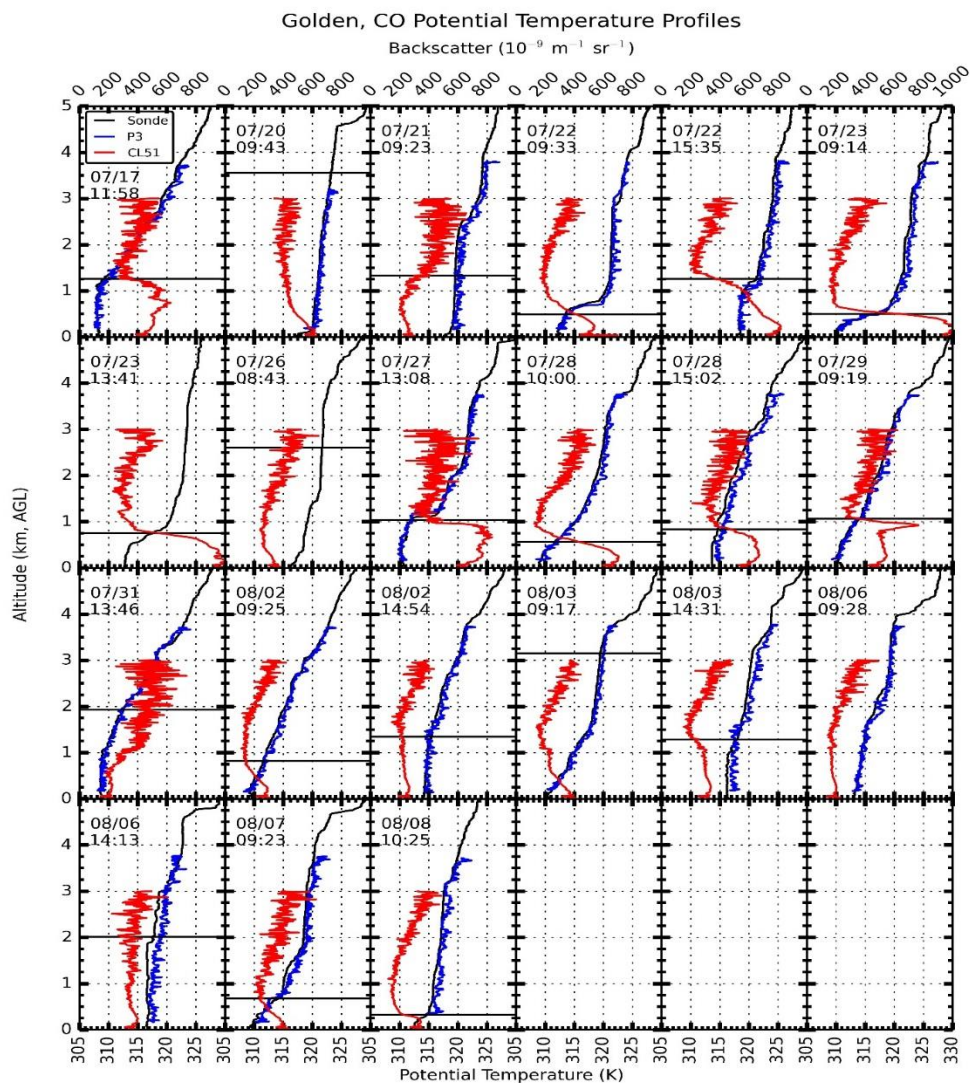
- Total of 53 radio sondes used for evaluation
- CL-51 data averaged over 5-minutes to account for spatial differences with sondes
- R = unfiltered
- R = filtered; 5-minute  $\sigma > 0.20$  km or RSD > 20%





# NASA P-3B Spirals used to Evaluate Spatial Representation of Sonde PBLH

- Potential temperature (**NASA P-3B** and **Millersville University radiosondes**) and **CL-51 backscatter** profiles collected at the Golden NREL site. Horizontal lines indicate MLH as determined via BLView.
- PBL Height from NASA P-3B spiral (~5 km) shows good agreement with sondes.





## Ad-hoc Ceilometer Evaluation Study (ACES)

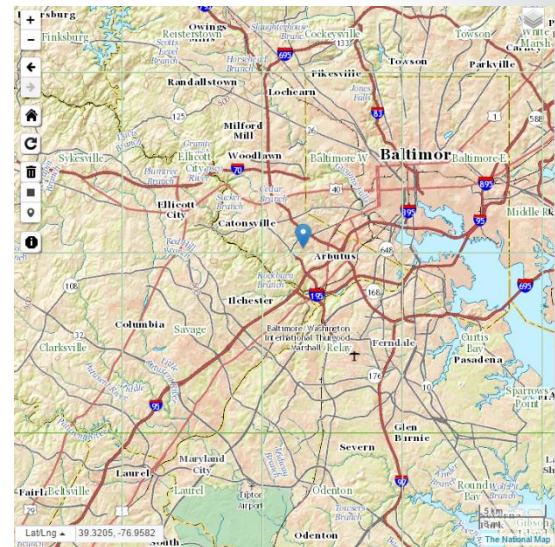
**Location: University Baltimore Maryland County -combines EPA PAMS evaluation effort with on-going ceilometer test-bed effort.**

**Main instruments - Vendor/Model: Campbell Scientific –CS135, Viasala – CL-31 and CL-51, and Lufft CHM15K.**

**Other relevant measurement: Sigma space Micro-pulse lidar, Leosphere ALS-450, and Radiometrics MWR, Leosphere windcube 200s**

**Study period: ~ November 15 through December 16, 2016+**

- CL-51, CS135, and CHM15K running since November 8
- Test ceilometer performance in low aerosol loading environment
- Assess the range corrected attenuated backscatter to identify aerosol layer heights for mixing layer height determination during morning and evening transition periods:
  - MLH using available vendor software
  - MLH using a common algorithm



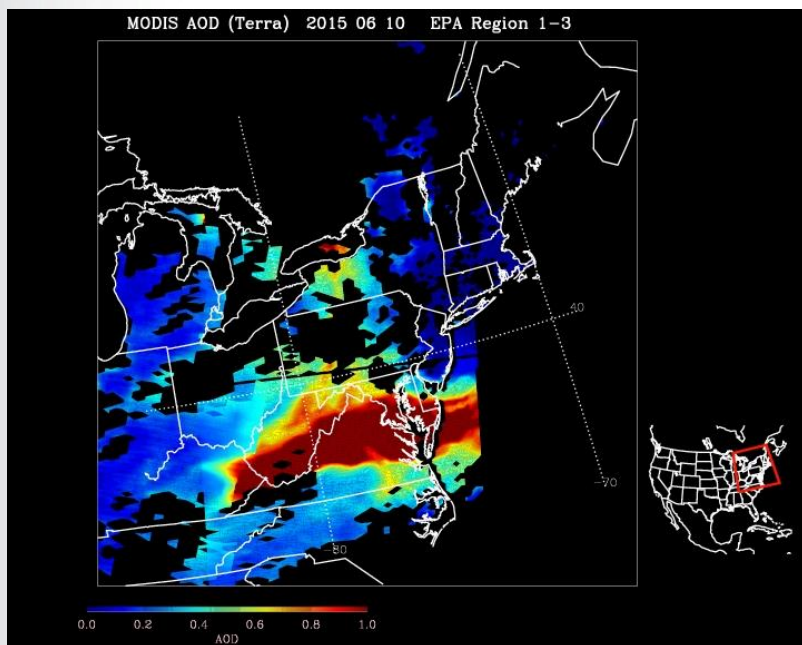


# Data Considerations/Thoughts

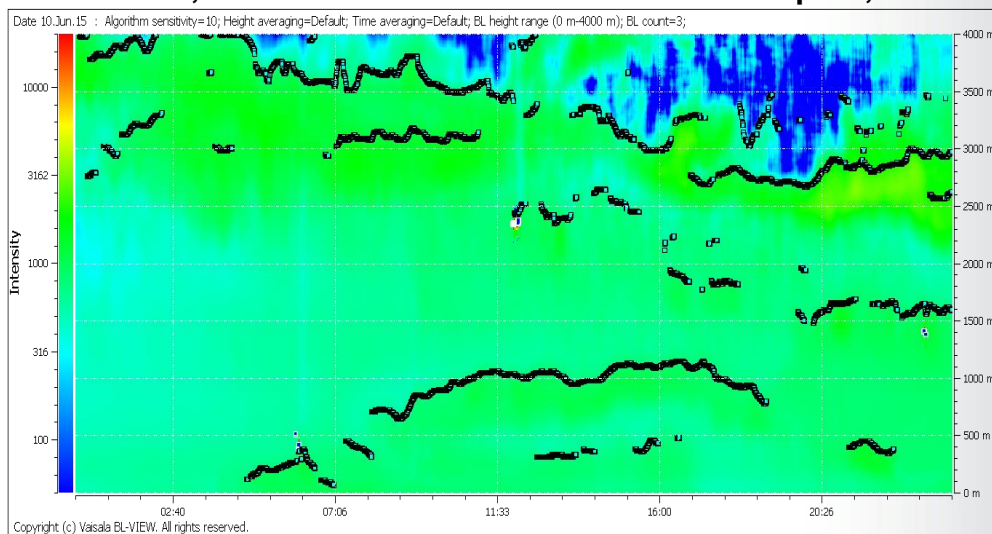
- Hourly Mixing Height is the required variable under PAMS
- Ceilometers capable of providing attenuated backscatter profiles up to 15km+
- Data logging the entire backscatter profile allows for:
  - Alternative algorithms to derive MLH, especially if other lidars (MPL, HSRL, CL-31s) are part of any larger network
  - Visual check on the derived MLH
  - Additional uses of data:
    - EPA exceptional events analysis
- Spatial variability of MLH around PAMS sites
- How to ensure a high quality MLH at PAMS sites operated by different state and local air quality agencies.
- Synergies with other networks



## June 10, 2015 – Canadian Forest Fires Smoke Plume



### Characteristic backscatter curtain plot generated in BL-View for 10-June, 2015 at CAPABLE Research Site - Hampton, VA



**Smoke from the Canadian forest fire was observed by increased backscatter in the 2500 – 4000 m range.**